

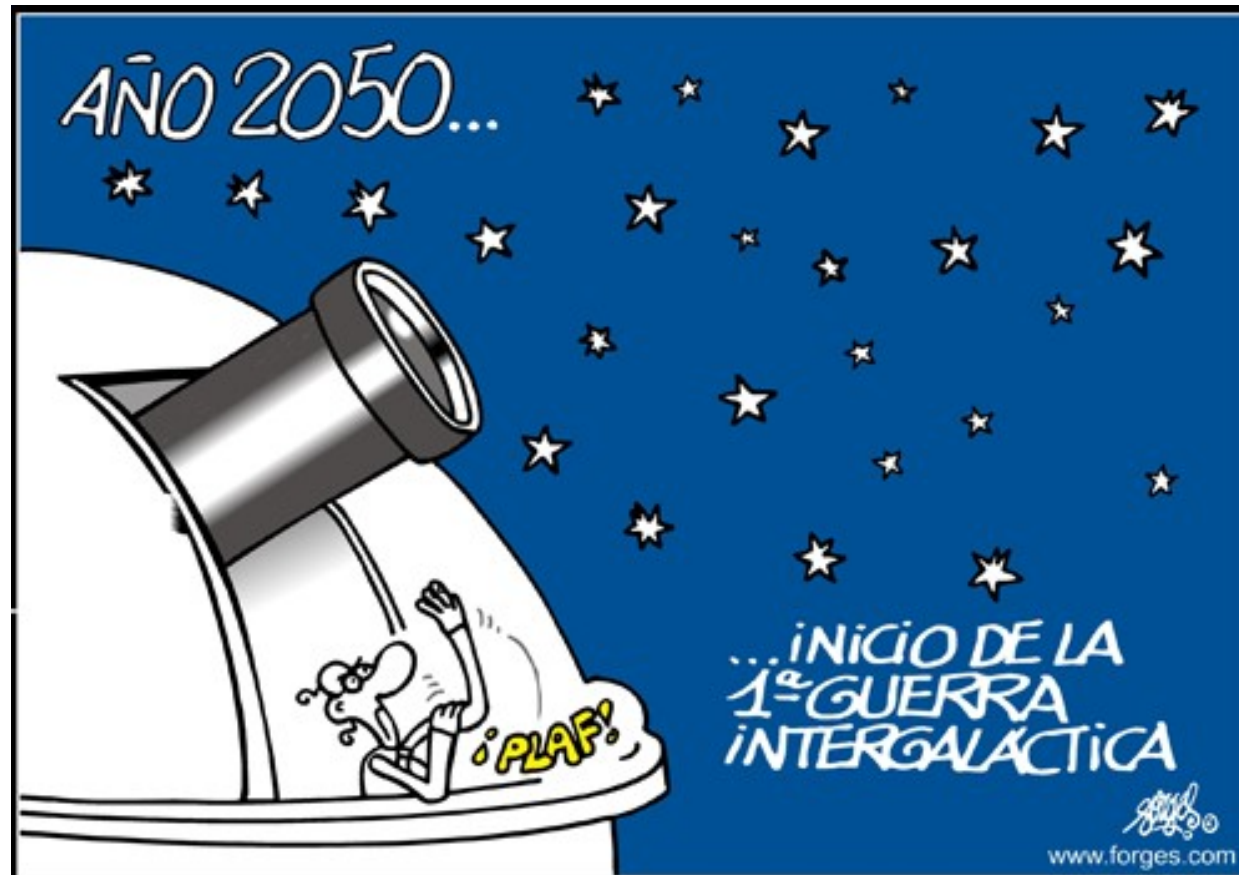
The Big Bang of astronomical data. How to use Python to survive the data flood.



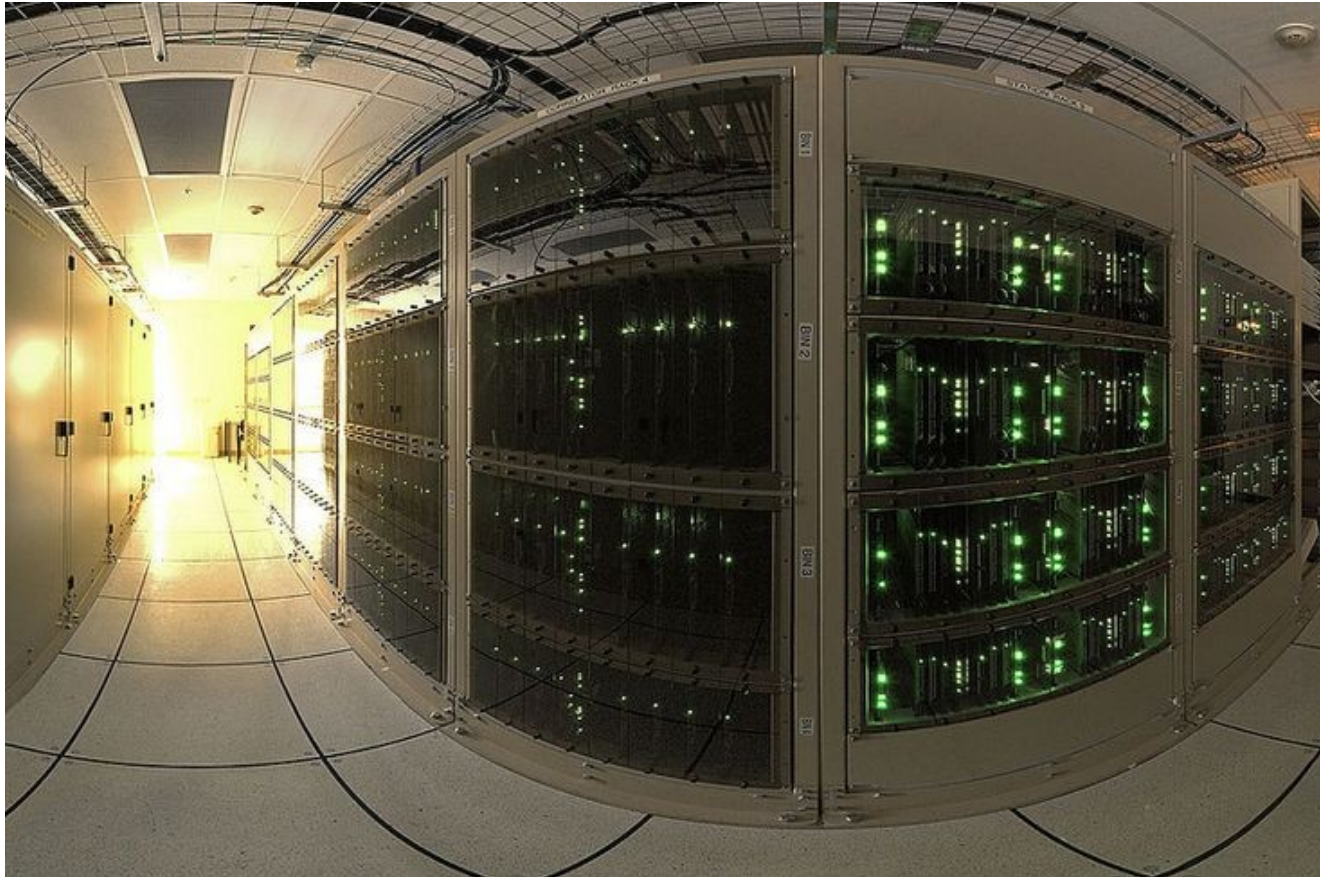
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J. Garrido, J. E. Ruiz, L. Verdes-Montenegro and
the LOFAR surveys team

The new astronomy



The new astronomy



ALMA correlator

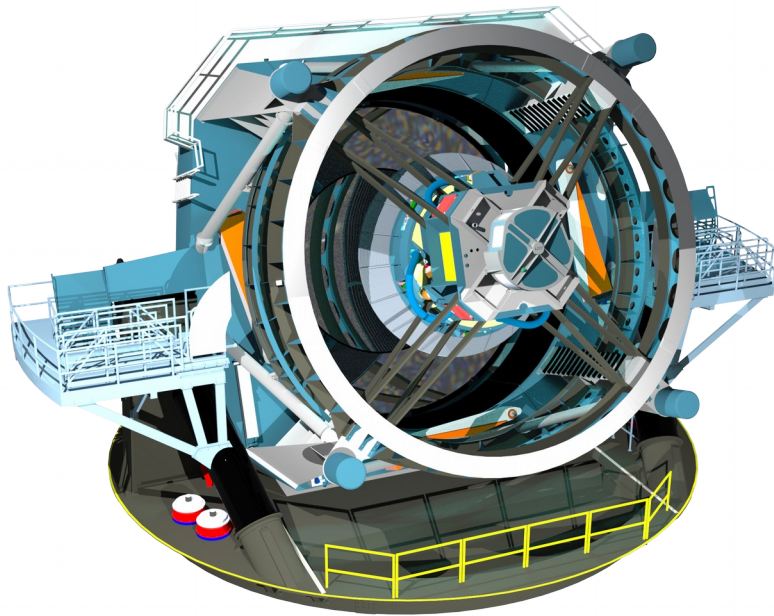
Astronomy and Python

- Python is currently the main language used for astronomy
- General Python computing libraries: numpy, scipy, matplotlib, pandas, emcee...
- Specific astronomical libraries (see <http://www.astropython.org/packages/>)
 - astroML: machine learning and data mining
 - astropy: main general library for astronomy
 - etc.

The future of astronomy

- New state of the art astronomical infrastructures that produce an overwhelming amount of data
- Examples:
 - ESA Gaia
 - Large Synoptic Survey Telescope
 - ESA Euclid
 - The Square Kilometre Array and its pathfinders (LOFAR, ASKAP, Meerkat...)
 - Etc.

Large Synoptic Survey Telescope



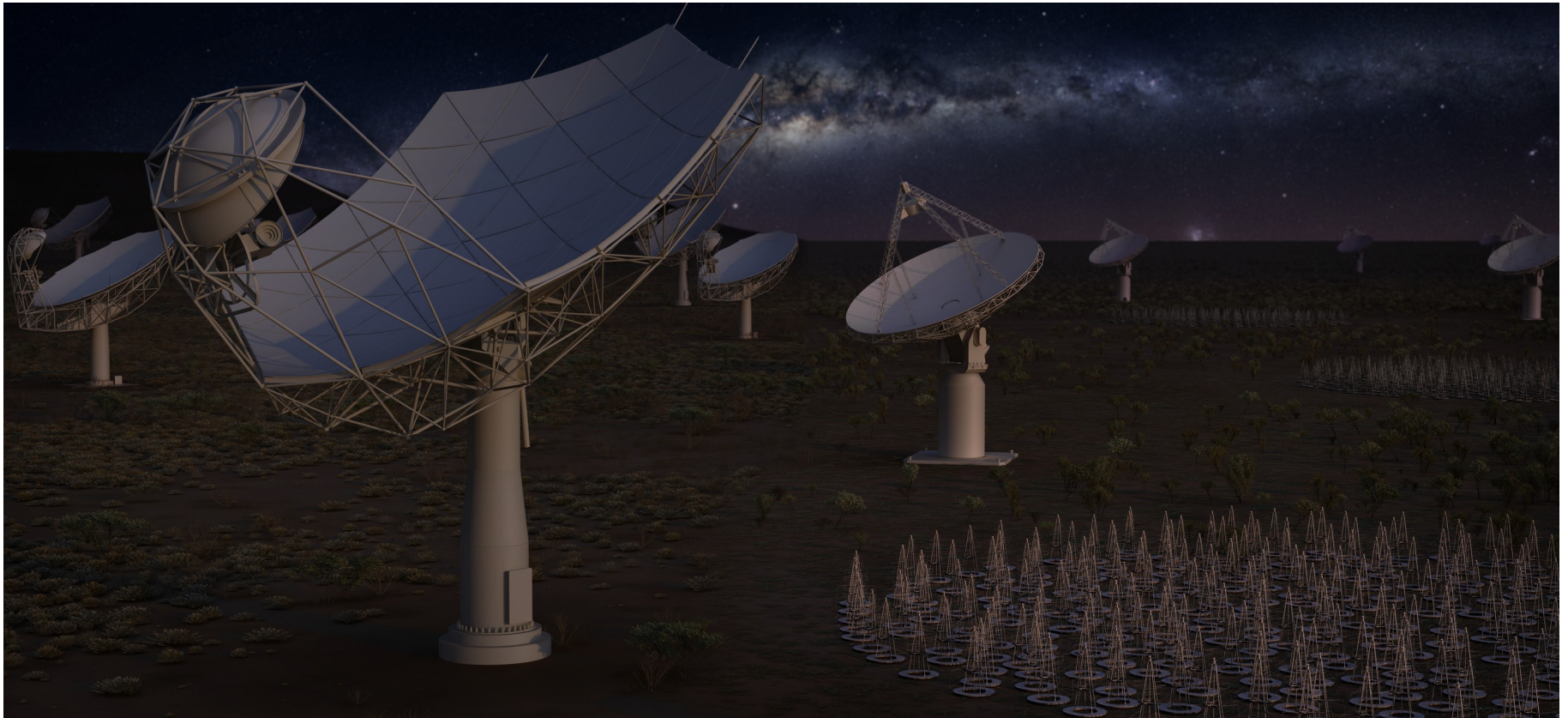
- 8.4 m mirror
- Covers the full visible sky every two nights
- Under construction - operational in 2022

Large Synoptic Survey Telescope



- Camera 189x16 Mpix
- Pipeline preprocessing: 3GB/s
- 30 TB per night during 10 years
- 2 M events triggered per night

Square Kilometre Array (SKA)



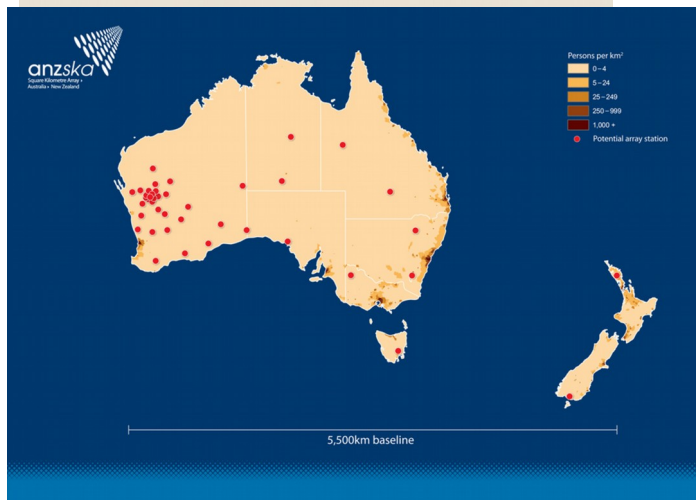
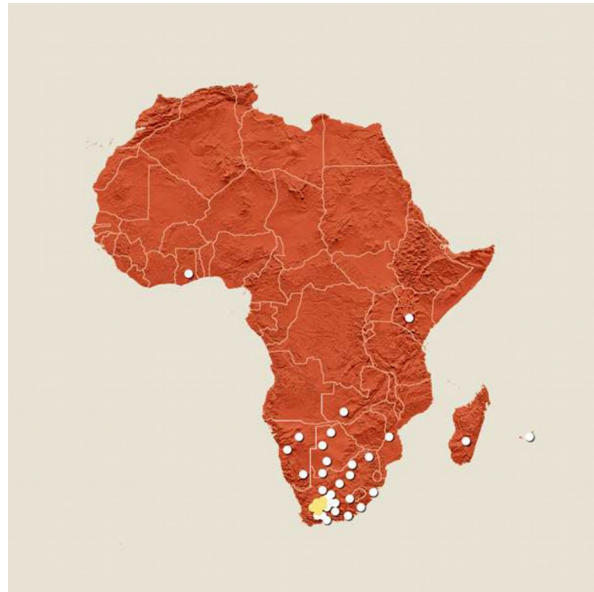
- Radio telescope with 1 km² of collecting area
- Phase 1 - 2020

SKA data



- Phase 1:
 - 10 TB/s from the antennas to the correlator
 - 40 GB/s of data → 70 PB per year
 - 1 MW infrastructure and 10 MW processing

SKA data



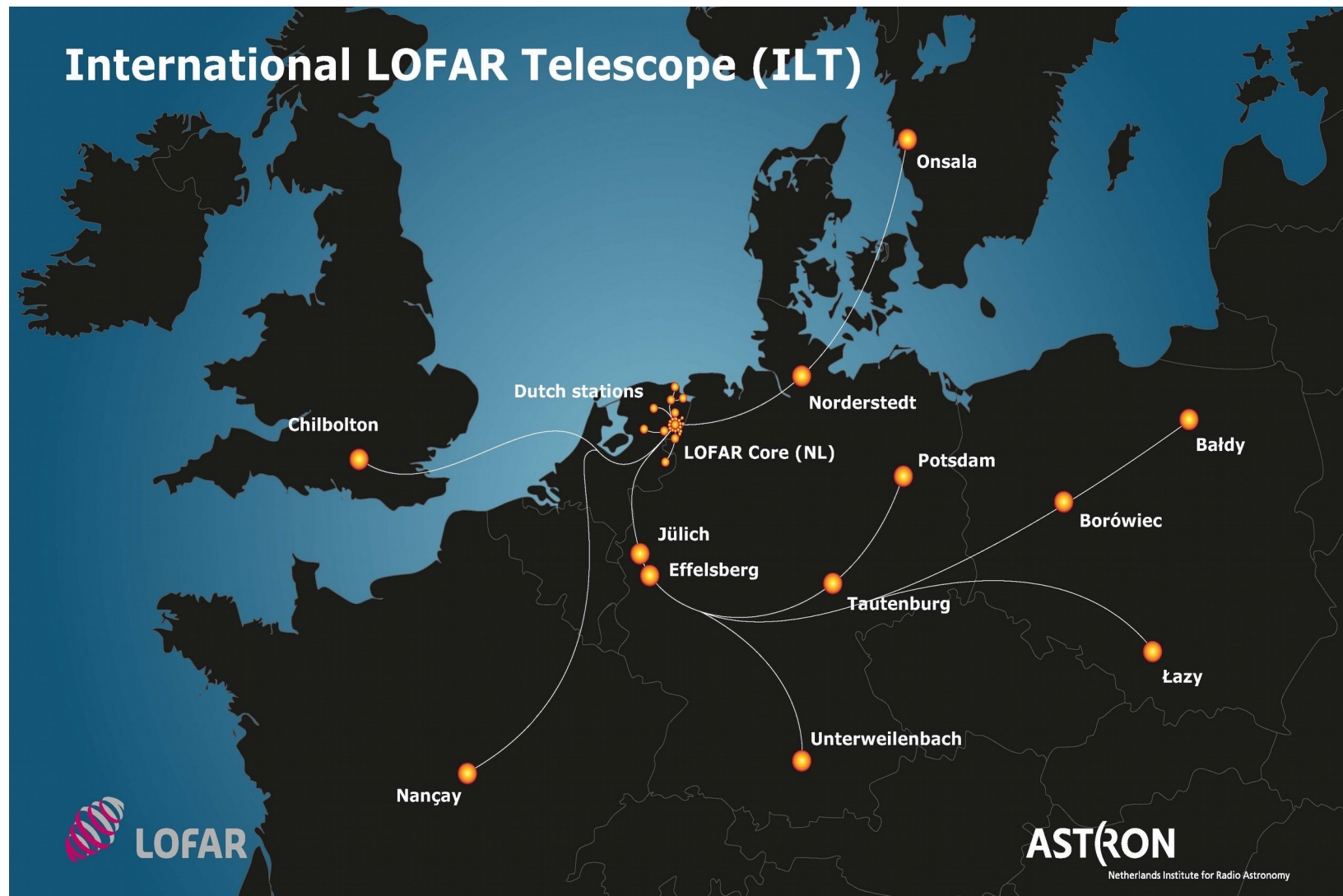
- Phase 2:
 - 160 TB/s from the antennas to the correlator
 - > 100 GB/s of data
→ 4.6 EB per year
 - 200 to 2000 dishes
 - 130K to 1M antennas

LOFAR

- Low Frequency Array
- Software defined radio-interferometer working at low frequencies (30 to 240 MHz)
- One of the Square Kilometre Array pathfinders



LOFAR Stations



LOFAR Stations



LOFAR frequencies

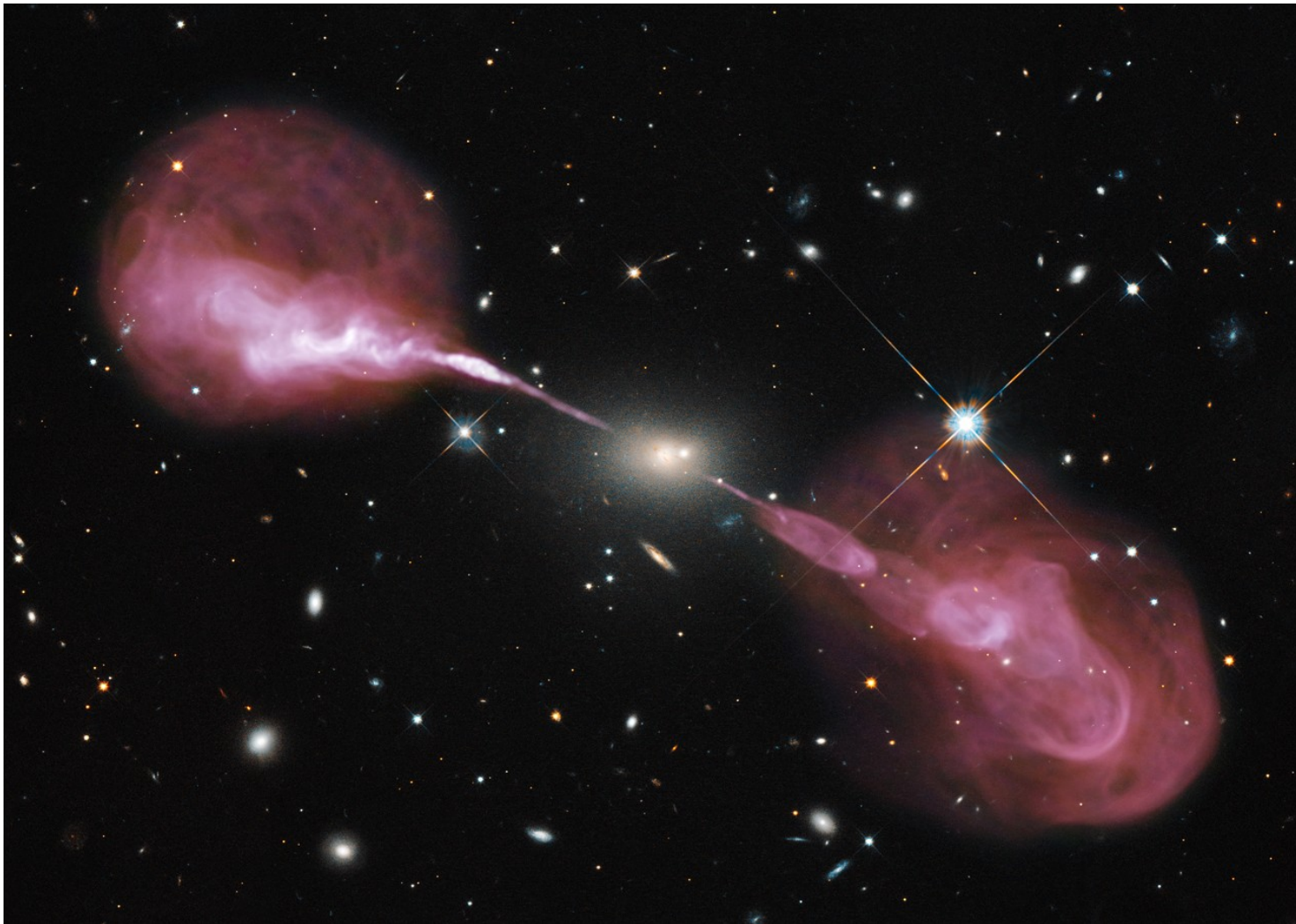
- LBA 30-80 MHz
- HBA 120-240 MHz



LOFAR science

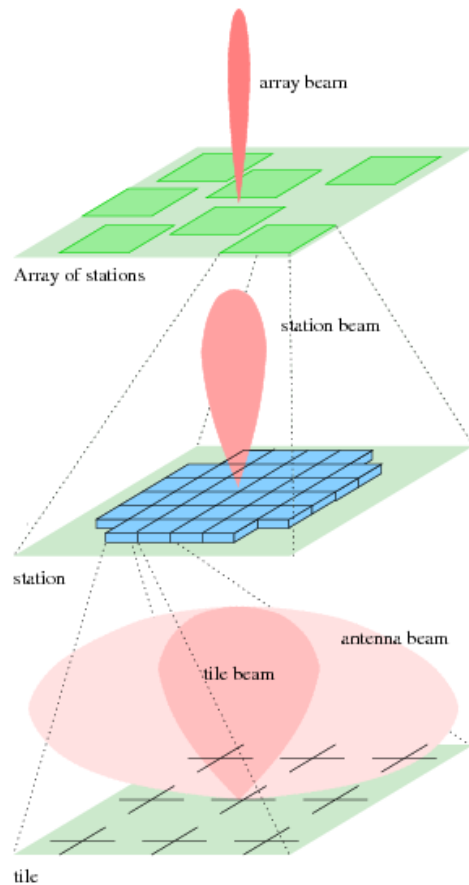
- Origin and evolution of galaxies and supermassive black holes
- Epoch of reionization
- Solar science and space weather
- Transients
- Map the galaxy using pulsars
- Exoplanets, SETI

Radio galaxies

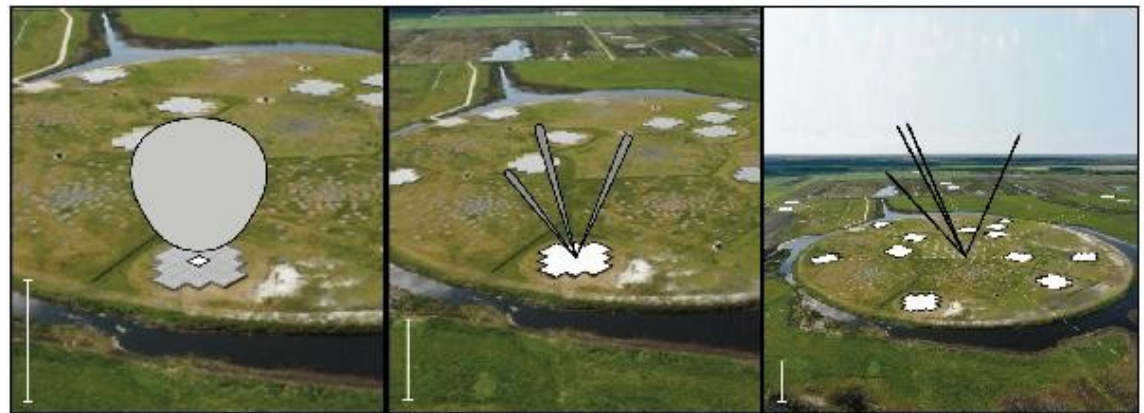


Hercules A. Credits: NASA and the NRAO

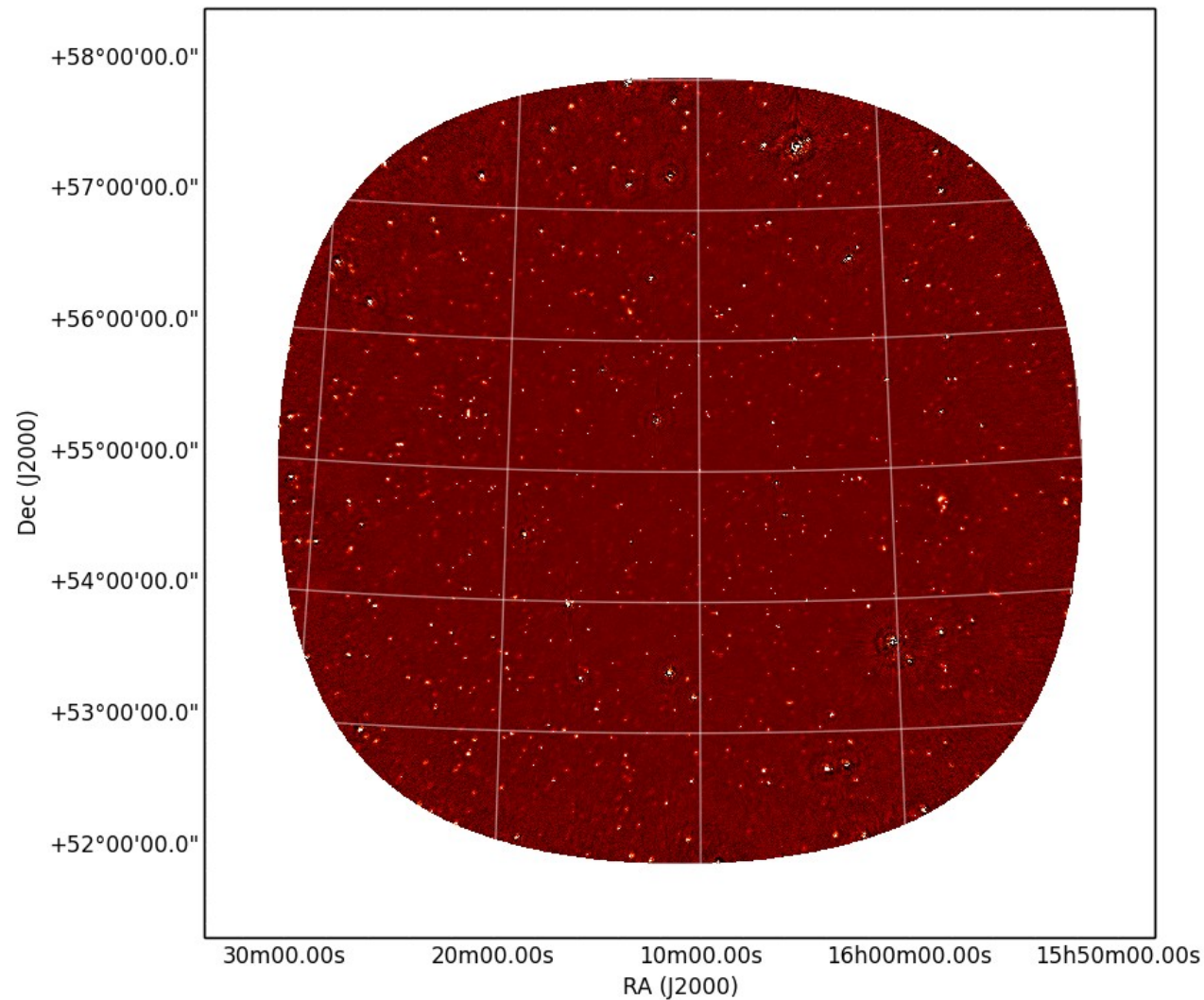
LOFAR aperture synthesis



- field of view diameter of ~ 5 deg at 150 MHz
- resolution < 5 arcsec (up to 0.1 arcsec)



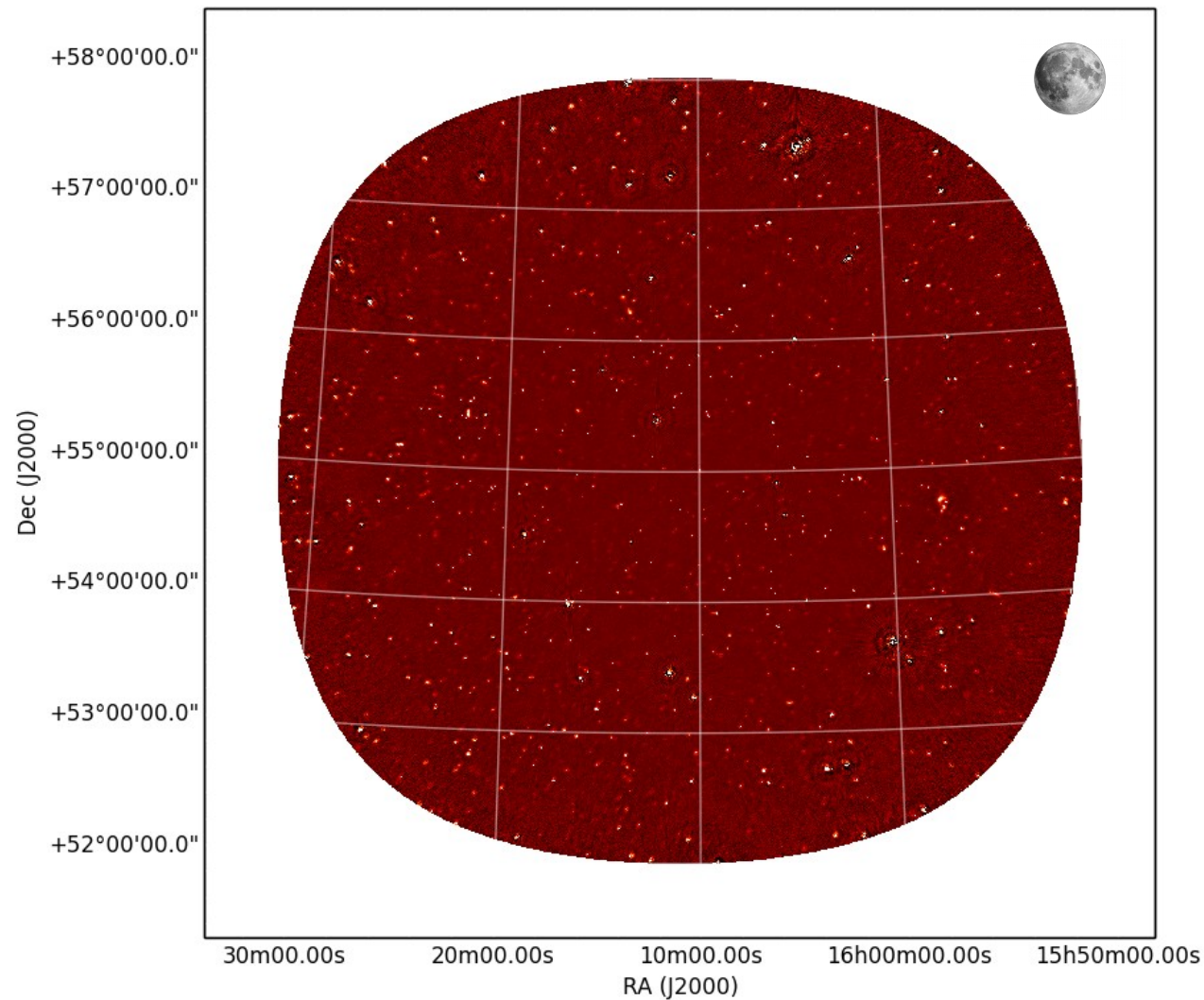
LOFAR imaging



In 8 hours
~40 sq. deg.
5000 sources

Calibration on
IAA (Granada) cluster

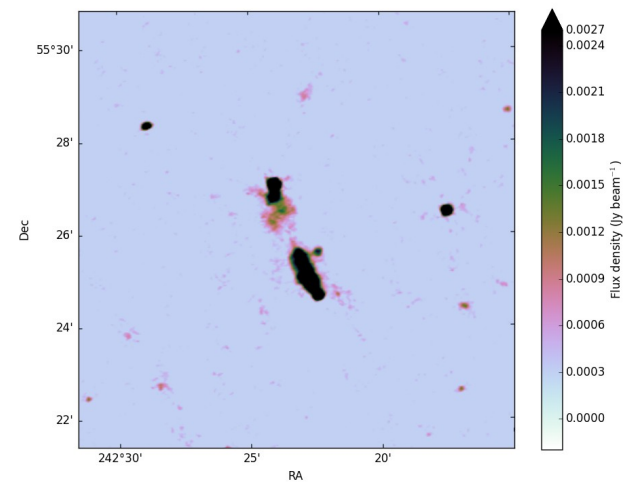
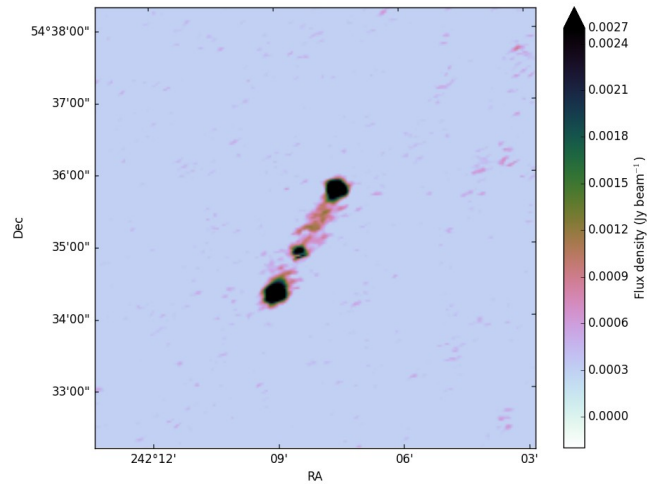
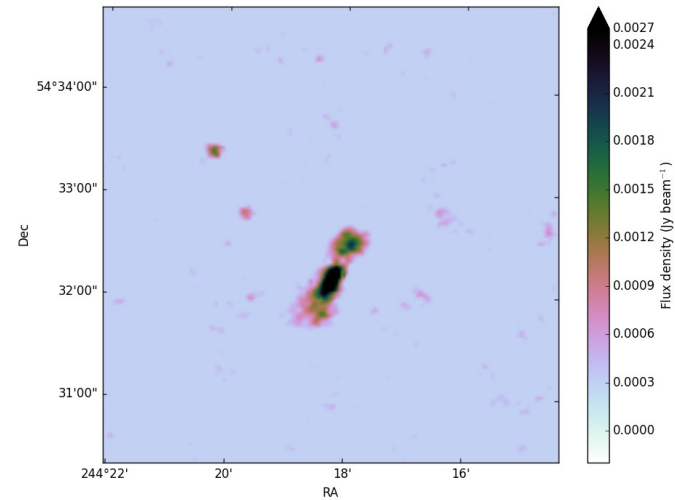
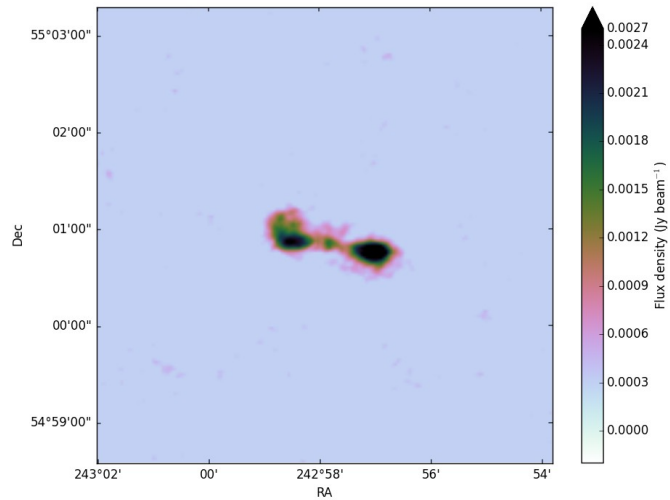
LOFAR imaging



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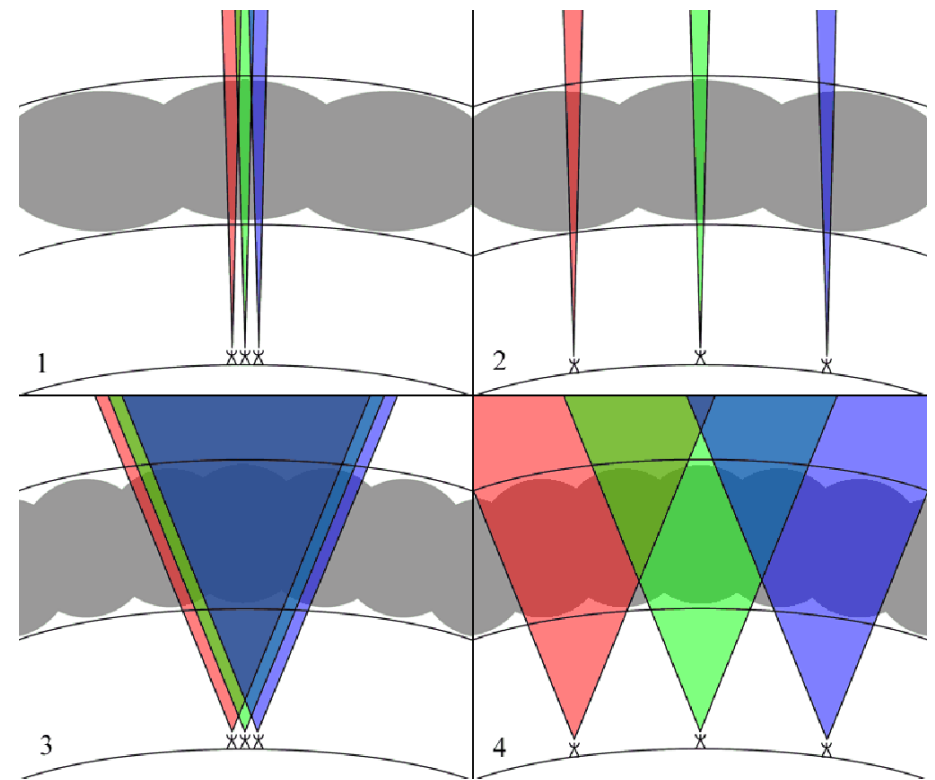
Calibration on
IAA (Granada) cluster

Extended sources



Ionosphere

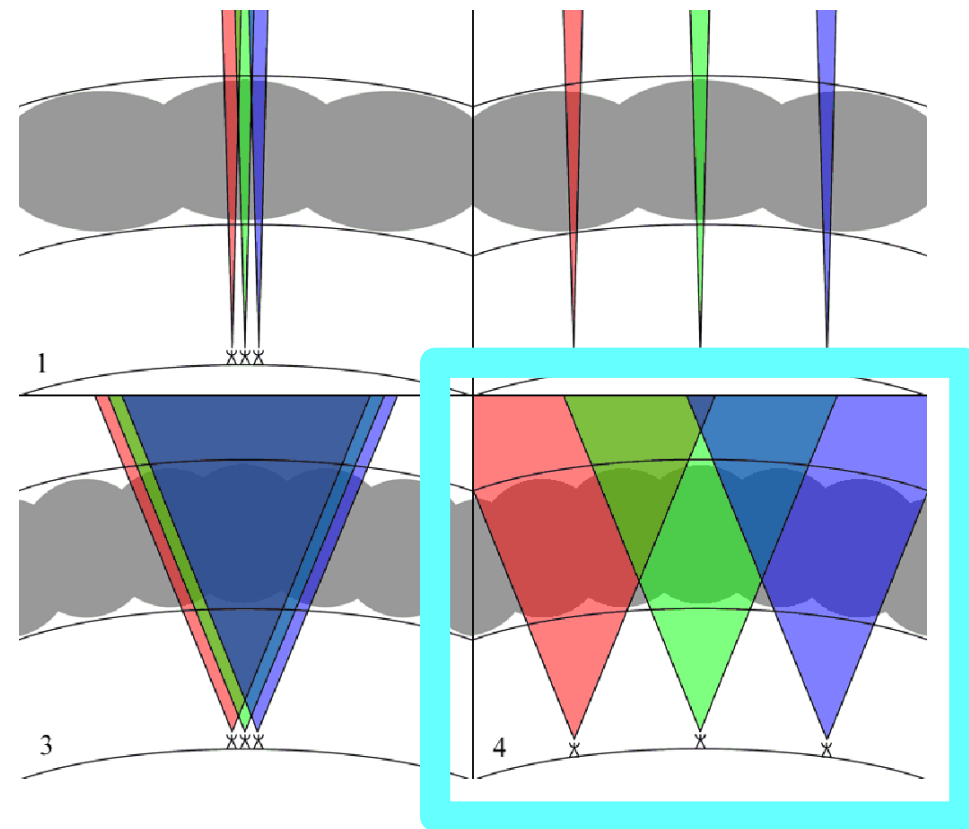
- Effect depends on frequency, length of the baselines and f.o.v.
- LOFAR, worst case:
 - Wide field of view
 - Long distance baselines
 - Low frequency



H. Intema

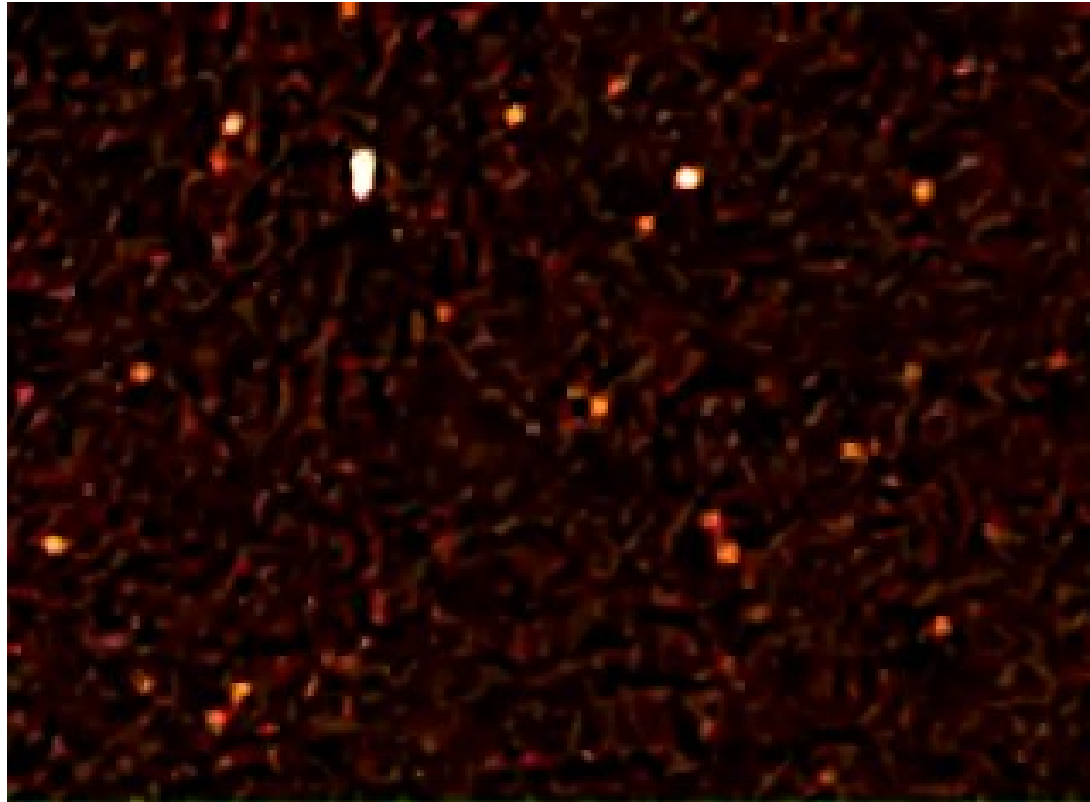
Ionosphere

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H. Intema

Ionosphere



Challenges for the astronomer

- User data calibration (remove the effect of the ionosphere and the RFI)
 - 8 hours full resolution → ~20 TB
 - Minimum of 2 CPU years to run the calibration
 - Experimental pipeline
- LOFAR calibration software
 - Difficult to install
 - Continuous development

Computational solution needed

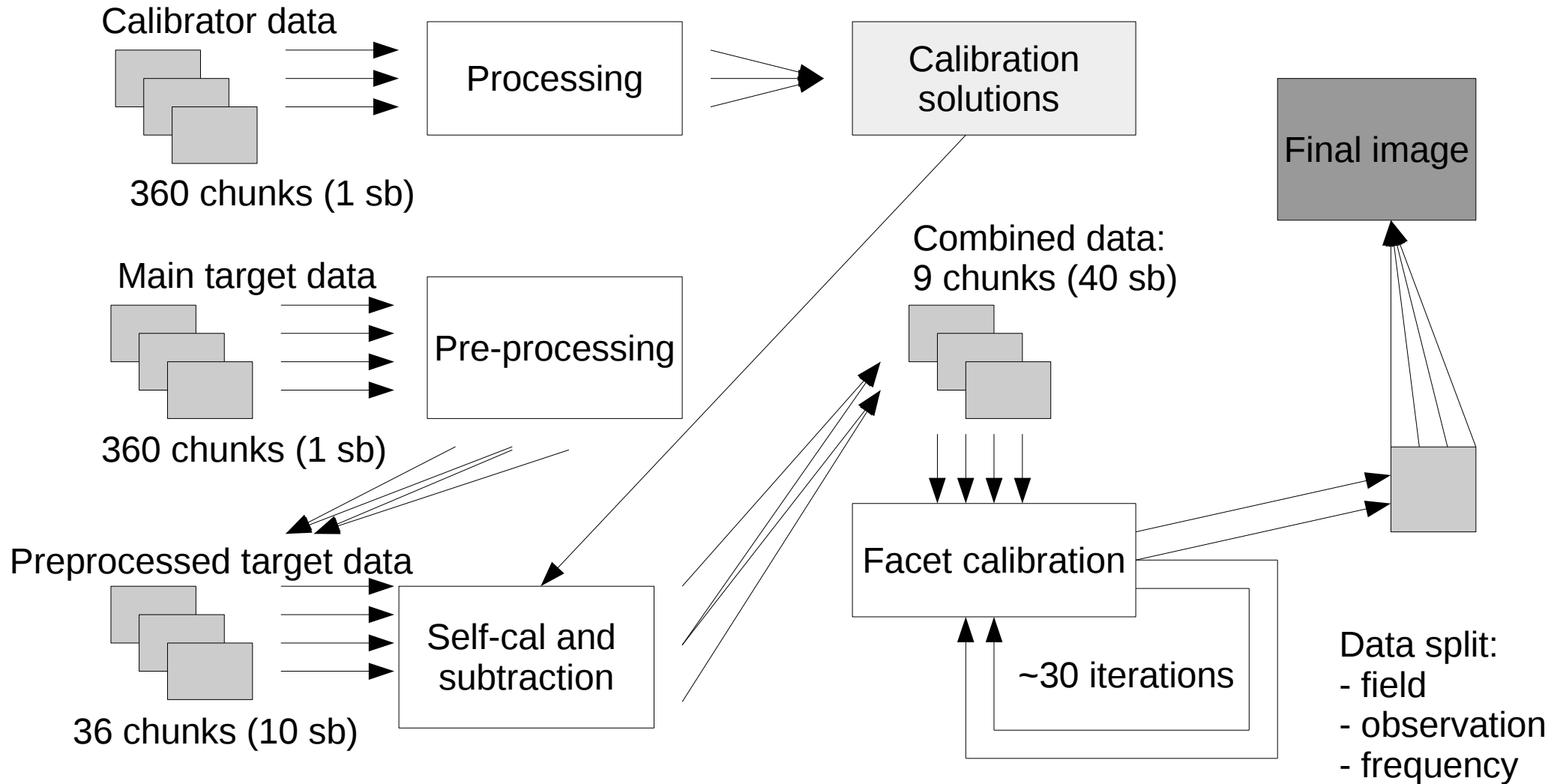
- Parallelizable:
 - Deal with a large amount of data in a reasonable time.
- Flexible:
 - Adapt the infrastructure (“hardware”) to different calibration strategies
 - Deal with quickly changing temperamental software
 - On-demand (optional but very useful)

HPC, HTC and cloud computing

- Tests in different infrastructures: clusters, GRID, cloud, etcetera.
- SKA-AWS astrocompute proposal
 - Preparation of the base infrastructure (virtual machine images, check provisioning of spot instances, etc.)
 - Data transfer: 50 TB
 - Adapt calibration pipeline and run

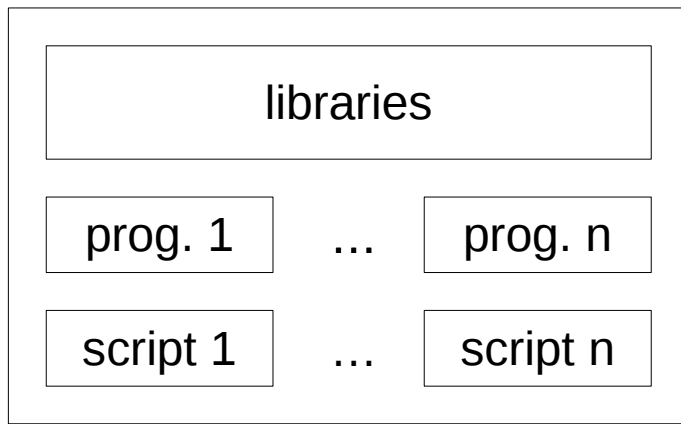
<http://www.lofarcloud.uk>

Experimental calibration pipeline



The role of Python

LOFAR software

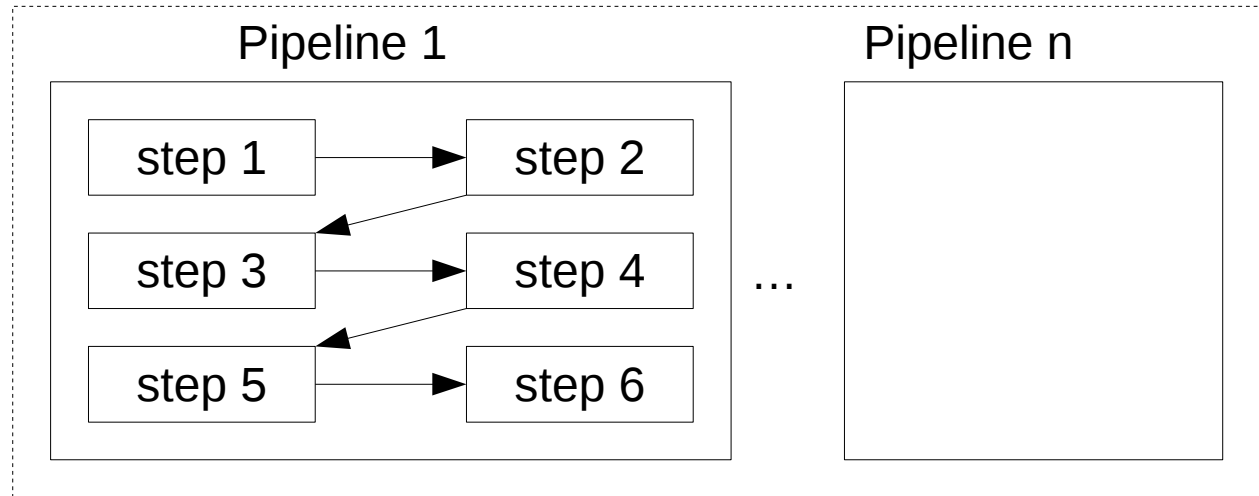


The role of Python

LOFAR software

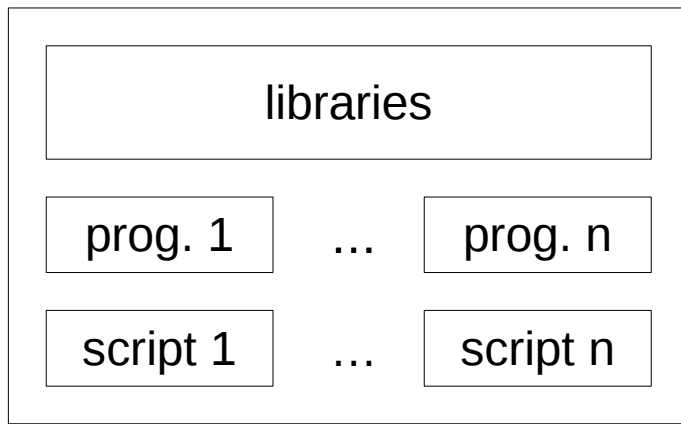


Experimental pipelines

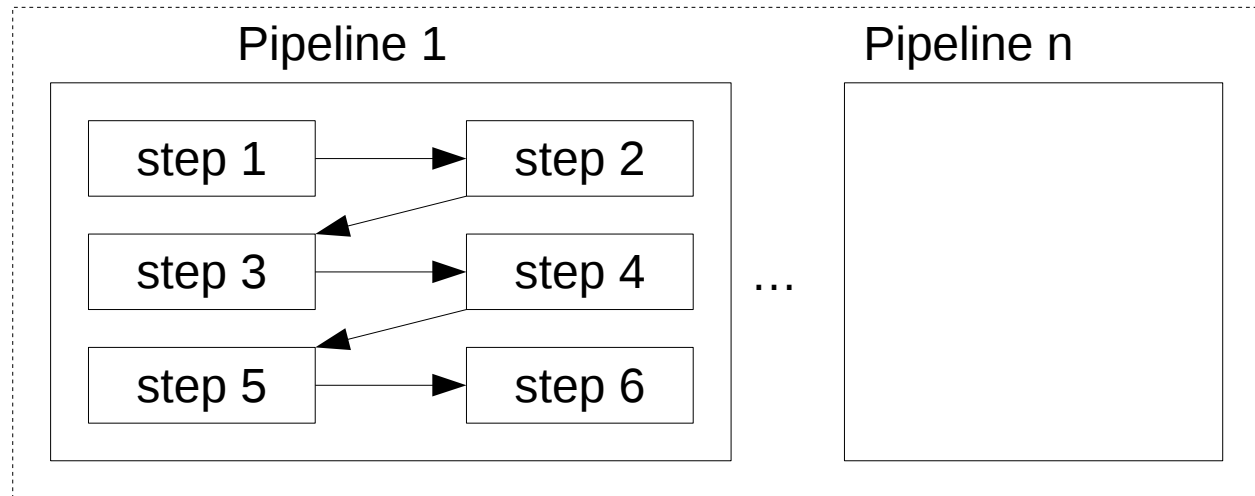


The role of Python

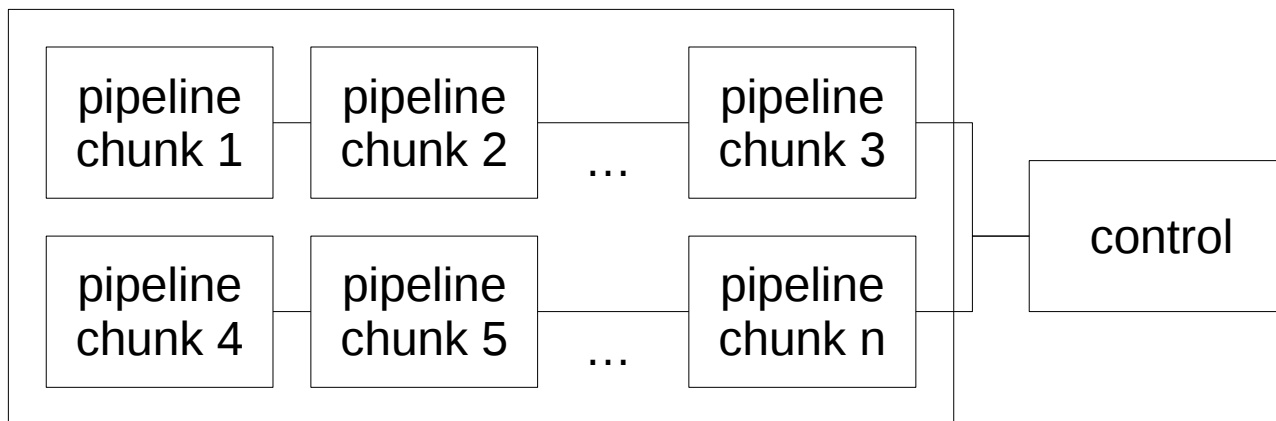
LOFAR software



Experimental pipelines

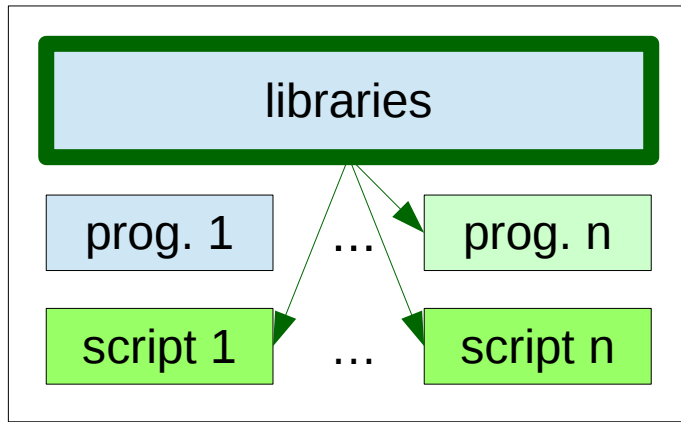


Infrastructure

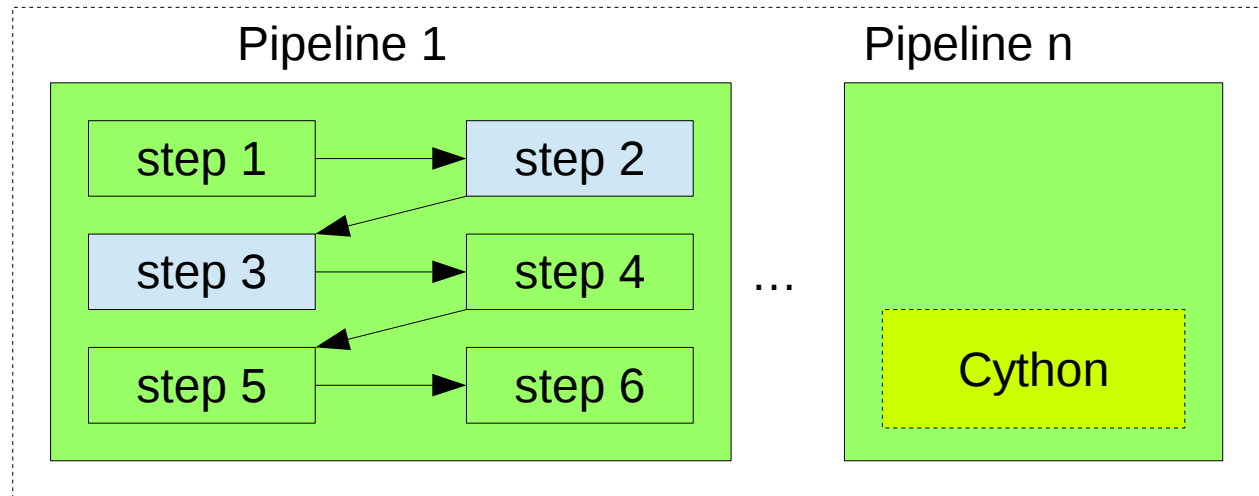


The role of Python

LOFAR software

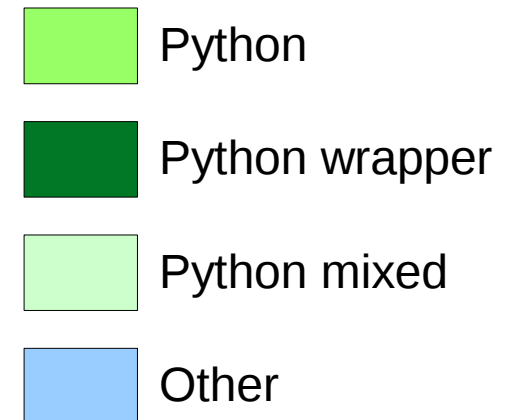
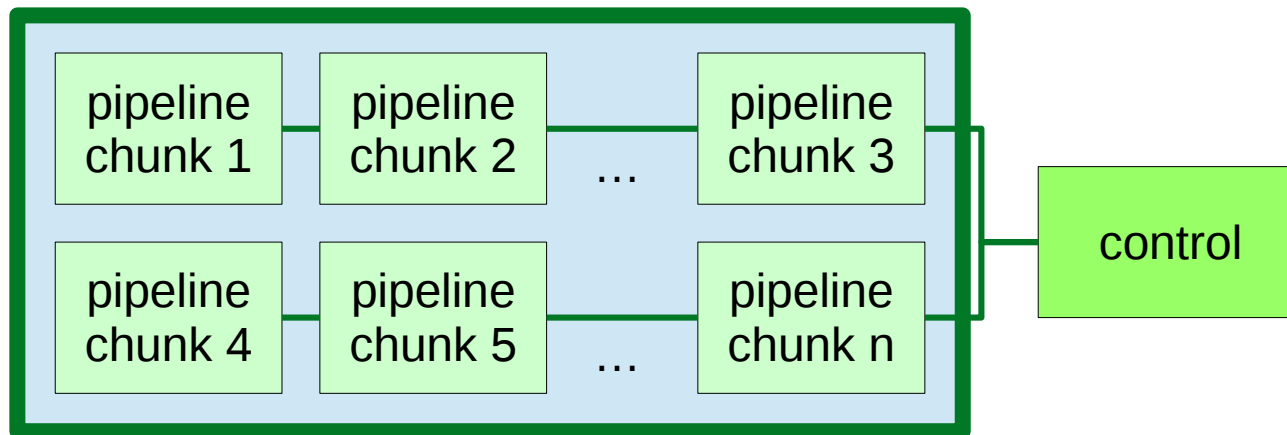


Experimental pipelines



Infrastructure

Ansible



Summary

- Big software and data managing challenges associated to new astronomical infrastructures, even for final users.
- The role of Python:
 - Quick prototyping - fundamental for experimental pipelines and testing.
 - Multi-domain - Can be used for a wide range of problems.
 - Robust - Enough to write “real” efficient software.
 - Unifying tool - that holds all together. 